



CTX and HiRISE 3D mapping using cloud computing

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Why process more DTMs of Mars

 Digital Terrain Models (DTMs) tell us about how the Martian surface was formed

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- Specifically, DTMs allow us to "read into the past" physical processes such as meteoritic impacts, volcanic lava flows, the action of water and dust and the motion of dunes.
- HRSC DTMs (from 50-150m/gridpoint) allow us to obtain a synoptic view of the surface
- CTX DTMs (≈18m/gridpoint) allow us to look at more of the details of the physical processes

• HiRISE DTMs (≈0.75m) and their time evolution) allow us to map the changes as they occur



Freie Universität Berlin Winted KINGDOM - CHINA - MALAYSIA OLR CASP-GO Auto-DTM processing chain

The University of



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Co-registered Ames Stereo Pipeline with Gotcha refinement and Optimization (CASP-GO) has improved performance compared to the original open source NASA Ames Stereo Pipeline system:

- (a) Co-registered geo-spatial coordinates wrt HRSC (and MOLA) data;
- (b) Improved DTM completeness for unmatched areas;
- (c) Reduced DTM artefacts;
- (d) Improved DTM accuracy;
- (e) Fully documented uncertainty value for every interpolated gridpoint





DTM production on MSSL's Imaging Cluster



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- Initial experiments based on the CTX & HiRISE stereo pairs for the 3 Mars rover sites.
- CASP-GO development based on QA from the 3 rover site data and then ported onto the MSSL's imaging processing blades for batch processing of MC11 CTX DTMs.
- A total of 117 CTX stereo pairs were processed for MC11 E (78) &W (39) on MSSL's imaging cluster.





Batch CTX DTM processing using Microsoft[®] Azure[®] cloud computing resources







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Examples of CASP-GO processed HiRISE DTMs

ESP_045062_1555_RED ESP_045128_1555_RED

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ESP_047185_1560_RED ESP_047119_1560_RED

